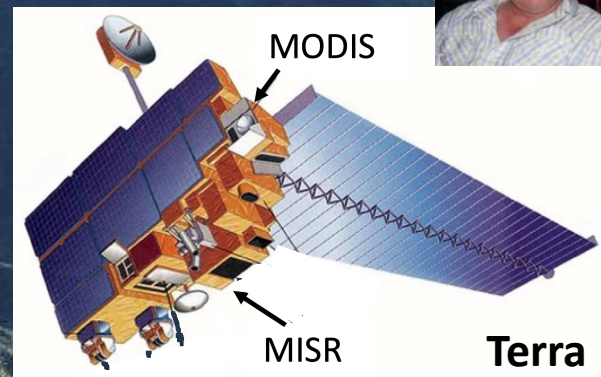
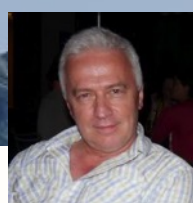


Understanding the Phenomenology of Opaque 3D Cloud Image Formation: Another Step Toward Cloud Tomography from Space-Based Imaging at Moderate Resolution



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IEEE International Conference on Computational Photography (ICCP) 2021

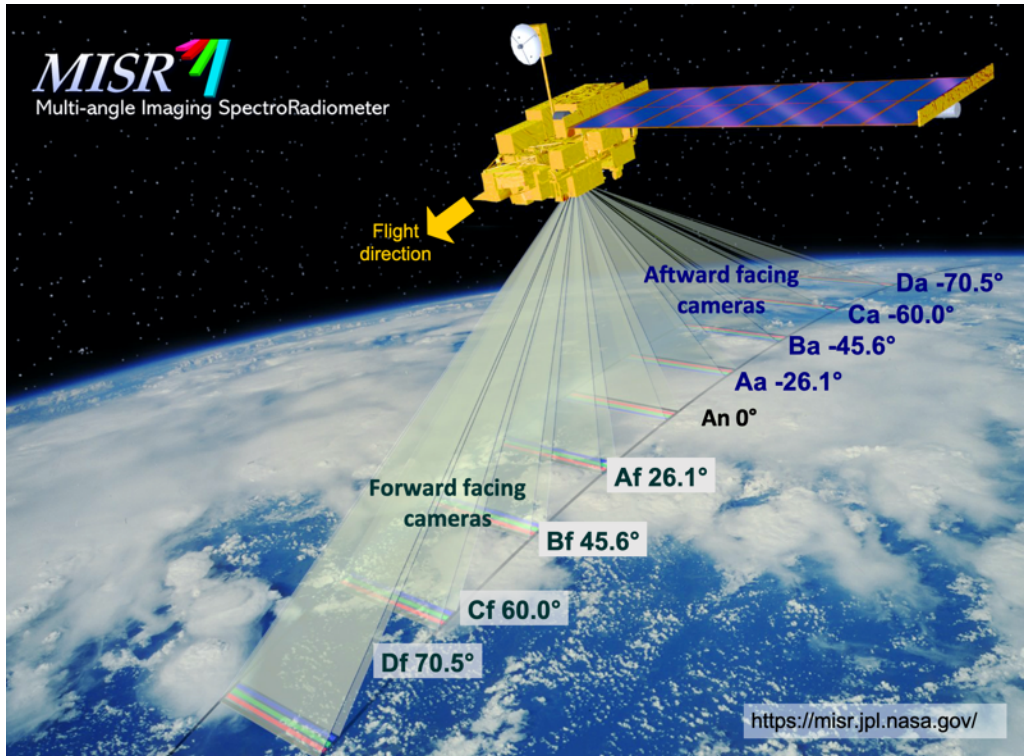
(2) Ludwig-Maximilians-University, Meteorological Institute



Physics & Optics Track



Multi-angle Imaging Spectro-Radiometer (MISR) / Terra (launched 1999)



Official cloud products:

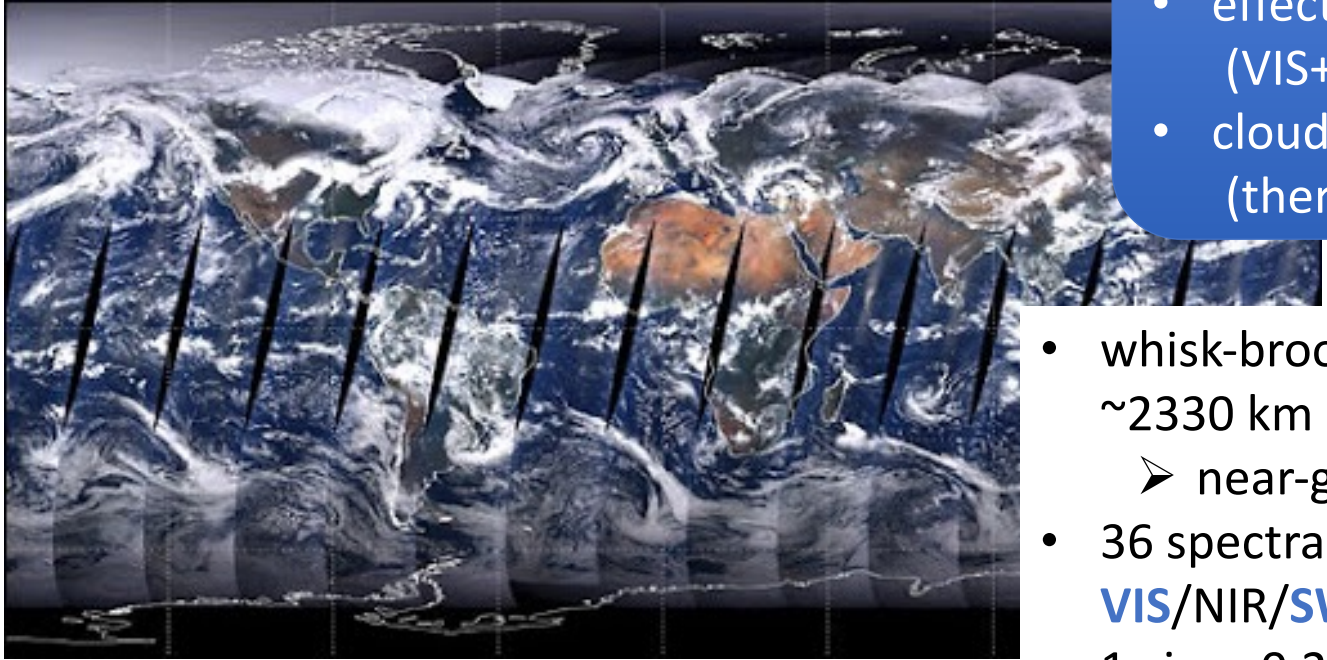
- cloud top heights
- height-resolved winds (stereo with time-delay)

- push-broom acquisition, ~400 km swath
 - global coverage in 9 days
- 4 spectral channels, all VNIR
- 9 views, 275 m pixels (always for the **red channel** used here) ≈7 minutes from most fore-ward to most aft-word images





MODerate-resolution Imaging Spectrometer (MODIS) / Terra (launched 1999)

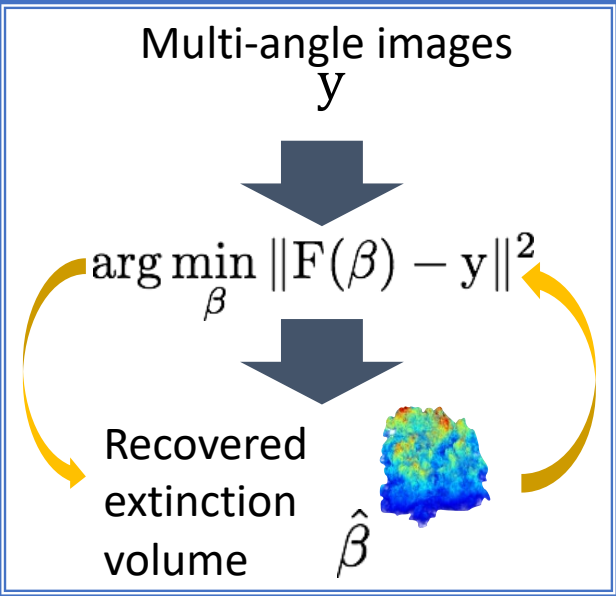


- Official cloud products:**
- cloud optical thickness
 - effective particle radius (VIS+SWIR algorithm)
 - cloud top height (thermal IR channels)

- whisk-broom acquisition, ~2330 km swath
 - near-global coverage every day
- 36 spectral channels, **VIS/NIR/SWIR/MWIR/LWIR**
- 1 view, 0.25–0.50–1.0 km pixels (as wavelength increases)

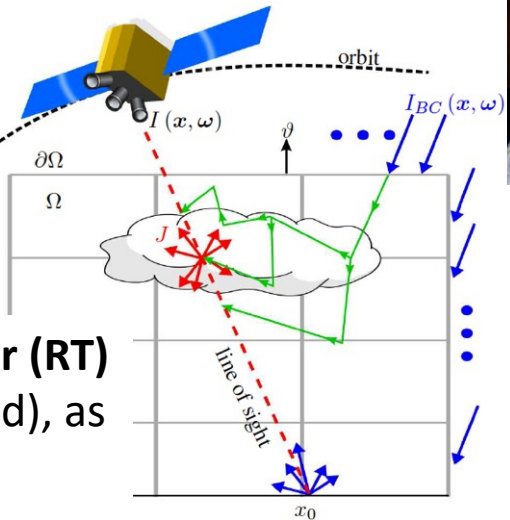


3D cloud tomography: Principles



“ β ” denotes a 3D gridded field of *unknown* extinction coefficient values.

Need a 3D radiative transfer (RT) solver: SHDOM (restructured), as forward model $F(\beta)$.



3D RT formulated as two *coupled* integral equations

↓ Formal solution of integro-differential 3D RT equation (a.k.a.

“upwind sweep”)

spatial integration along beam $(x, \omega) \rightarrow$

$$I(x, \omega) = \int J(x', \omega) \beta(x') e^{-\int \beta(r) dr} dx' + I_{BC} e^{-\int \beta(x) dx'}$$

Beer's law

propagation

directional integration over incoming $\omega' \rightarrow$

$$J(x, \omega) = \frac{\omega}{4\pi} \int_{s^2} p(x, \omega \cdot \omega') I(x, \omega') d\omega'$$

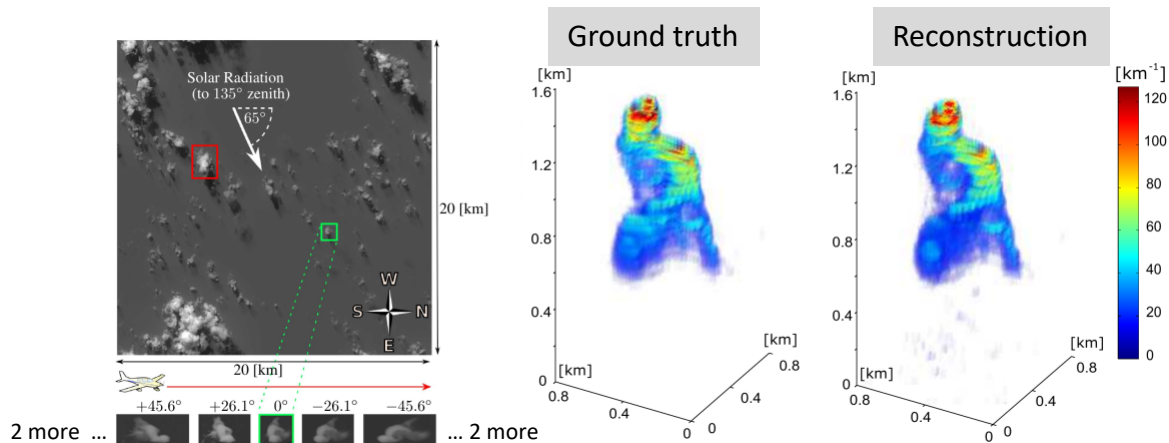
solar source term

scattering

↑ Definition of source function $J(x, \omega)$



3D cloud tomography: *Demonstration*



→ Progress toward
CloudCT mission
success:
poster presentations
#12, #16, #17 & #91

- Levis et al. (2015): red channel only, known microphysics (r_e, v_e), 9 views, **20 m resolution**
 - 46,656 unknowns & 315,018 unknowns, 2-step iteration scheme (1st being linearized) using SHDOM
 - application to data from **Airborne Multi-angle Imaging SpectroPolarimeter (AirMSPI)**, *also 20 m resolution*
- Levis et al. (2017): VNIR multi-spectral
 - basic (profile-only) microphysics (r_e, v_e) **without SWIR (à la MODIS) nor polarization (à la POLDER)**
- Levis et al. (2020): VNIR multi-spectral/multi-polarimetric
 - potential for a 3D full microphysics (N_e, r_e, v_e) retrieval using **polarization**: $[I, Q, U]$ Stokes vector components



Problem: airborne sensors have ≈ 20 m pixels

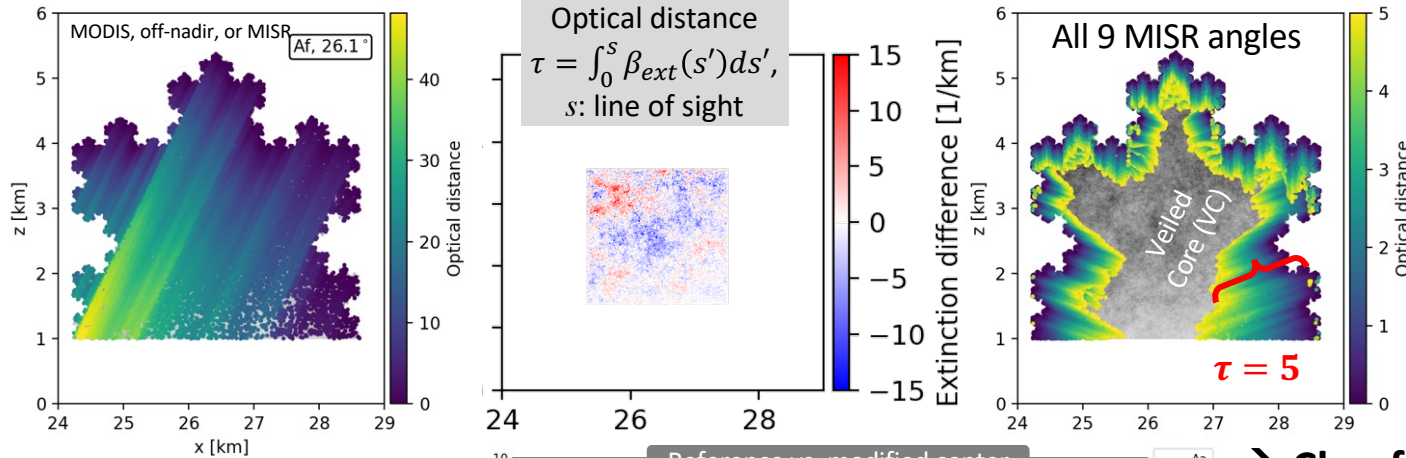
... while space-based ones (MISR + MODIS) have ≈ 250 & 500 m pixels!

→ *forward 3D RT modeling issues:* voxels can be opaque and/or internally variable

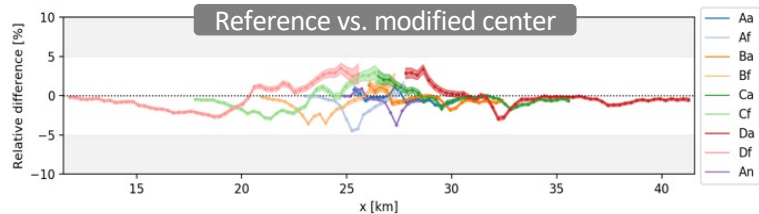
→ *inverse problem solution issues:* larger and more opaque clouds



The “veiled core” of opaque clouds



Manipulation of extinction field in the VC is *not* detectable in multi-angle imagery (lost in noise).



→ **Clue for *inverse problem* formulation:**
Just one-or-two unknowns in VC

Problem: airborne sensors have ≈ 20 m pixels

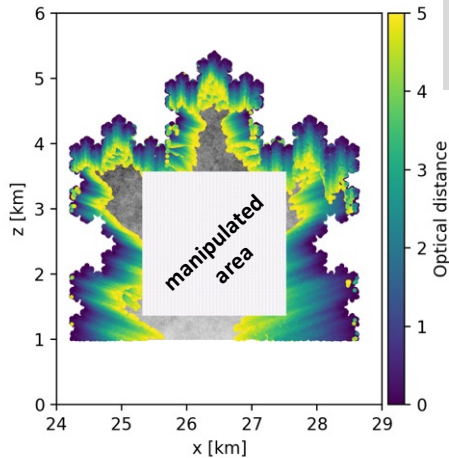
... while space-based ones (MISR + MODIS) have ≈ 250 & 500 m pixels!

→ *forward 3D RT modeling issues:* voxels can be opaque and/or internally variable

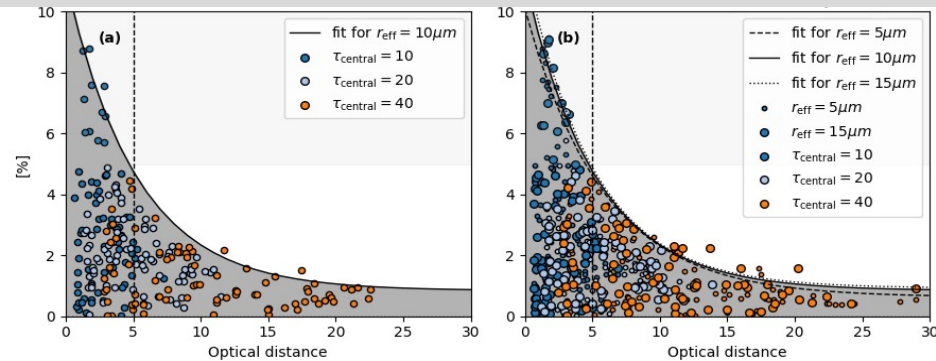
→ *inverse problem solution issues:* larger and more opaque clouds



The “veiled core” of opaque clouds



The $\tau \approx 5$ for “VC” threshold for $\approx 5\%$ tolerance is robust for clouds with sufficient opacity (say, maximum optical thickness in excess of ≈ 20).



L. Forster, A. B. Davis, B. Mayer, and D. J. Diner (2020), Toward Cloud Tomography from Space using MISR and MODIS: Locating the “Veiled Core” in Opaque Convective 3D Clouds, *J. Atmos. Sci.*, **78**, 155-166 (2021). DOI: <https://doi.org/10.1175/JAS-D-19-0262.1>



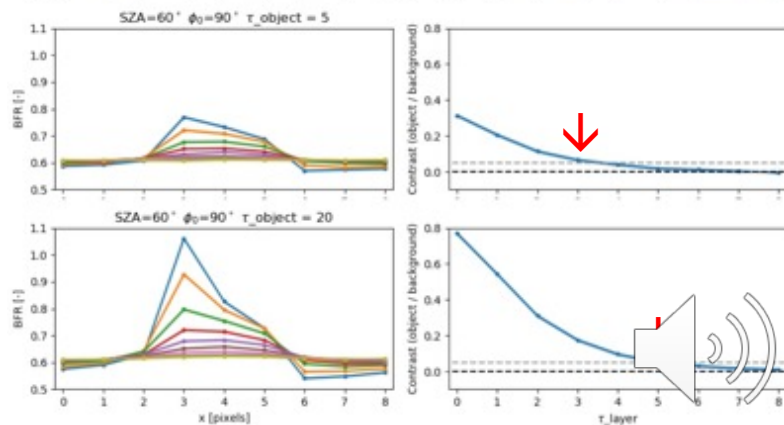
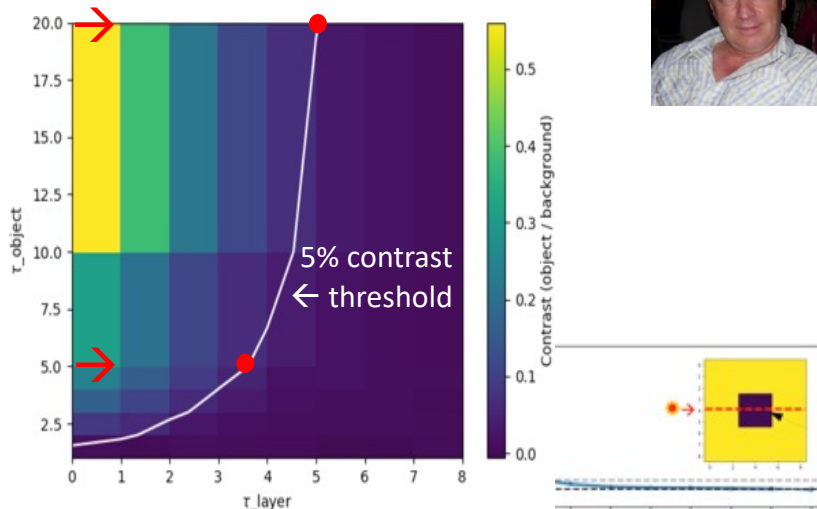
Cloud image formation in VNIR+SWIR: A tale of two diffusion processes



Diffusion process #1 & #2[†]

- random walks unfold on the unit sphere (i.e., **direction** space)
- in the **outer shell** (OS)
 - along-beam *drift* & lateral *dispersion*
- gradual loss of **directional** memory
- *pixel-scale* details in OS matter
- results in identifiable “**features**” in cloud imagery
- **RT regime:**
 - extinction and Beer’s law
 - forward-peaked scattering
 - small-angle/Fokker-Planck approximation

Superscripts “[†]” mean “adjoint” or “reciprocal” light, starting at the pixel/direction of interest in the image, propagating back into the cloud, and ending at sources.





Cloud image formation in VNIR+SWIR: *A tale of two diffusion processes*



Diffusion process #1 & #2[†] [or #1 & #3[†]]

- random walks unfold on the unit sphere (i.e., **direction** space)
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- **RT regime:**
 - extinction and Beer’s law
 - forward-peaked scattering
 - small-angle/Fokker-Planck approximation

Diffusion process [#2]

- random walks unfold in 3D **physical** space
- in the **veiled core** (VC)
- gradual loss of **positional** memory
- *cloud-scale* gradients in VC matter
- controls “**contrast**” between sunny and shaded sides
- **RT regime:**
 - scaled/transport extinction
 - effective isotropic scattering
 - diffusion/P₁ approximation



A. B. Davis, L. Forster, D. J. Diner, and B. Mayer (2020), Toward Cloud Tomography from Space using MISR and MODIS: The Physics of Image Formation for Opaque Convective Clouds, *J. Atmos. Sci.* (in preparation, preprint at <https://arxiv.org/abs/2011.14537>).





Cloud image formation in VNIR+SWIR: A tale of two diffusion processes



Diffusion process #1 & #2[†] [or #1 & #3[†]]:

- random walks unfold on 2D sphere (*direction space*)
- in the *outer shell*
- gradual loss of *directional* memory

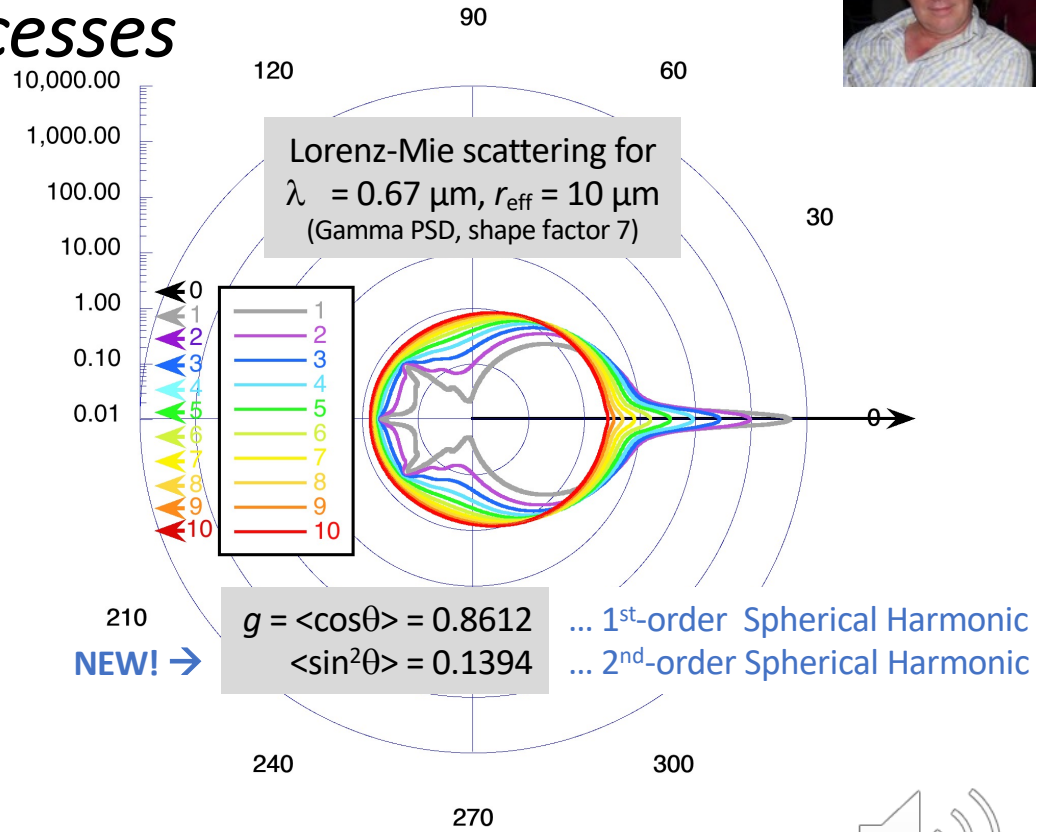
Characteristic (discrete) time scale to forget solar/sensor direction:

$$N^* = 1/\ln(1/g)$$

$$N^* \approx 6.6 \text{ for } g = 0.86$$

Associated with ...

- longitudinal drift: $\langle z_{N^*} \rangle = \ell(1-g^{N^*+1})/(1-g)$
- lateral dispersion: $2\langle \Delta x^2_{N^*} \rangle^{1/2}$



... explains empirical threshold ≈ 5 in optical distance that defines the VC.





Cloud image formation in VNIR+SWIR: A tale of two diffusion processes



Diffusion process #1 & #2[†] [or #1 & #3[†]]:

- random walks unfold on 2D sphere
(*direction space*)
- in the *outer shell*
- gradual loss of *directional* memory

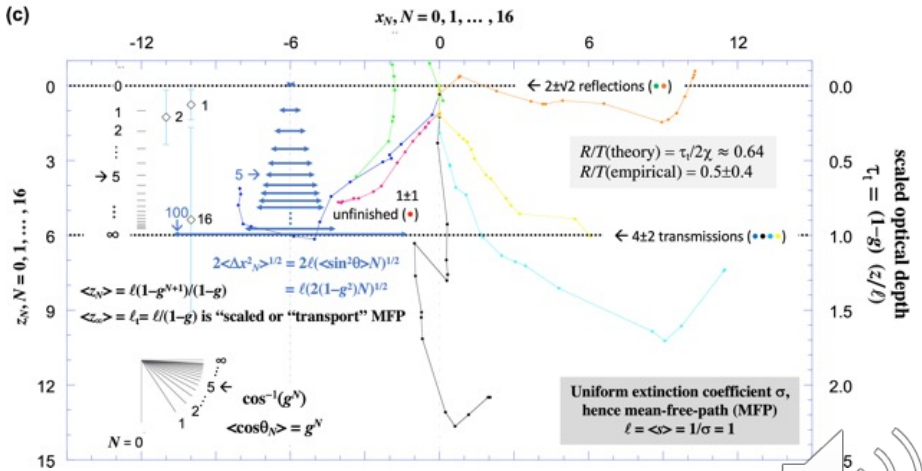
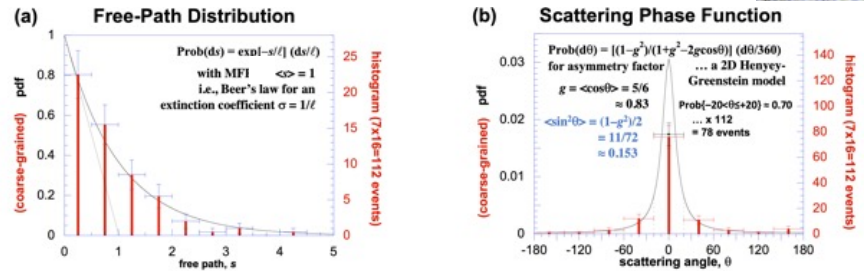
Characteristic (discrete) time scale
to forget solar/sensor direction:

$$N^* = 1/\ln(1/g) \approx (1/g - 1)^{-1}$$

$$N^* \approx 5 \text{ for } g = 5/6 \approx 0.83$$

Associated with ...

- longitudinal drift: $\langle z_{N^*} \rangle = \ell(1 - g^{N^*+1}) / (1 - g)$
- lateral dispersion: $2\langle \Delta x^2_{N^*} \rangle^{1/2}$



... explains empirical threshold ≈ 5 in optical distance that defines the VC.



Cloud image formation in VNIR+SWIR: A tale of two diffusion processes

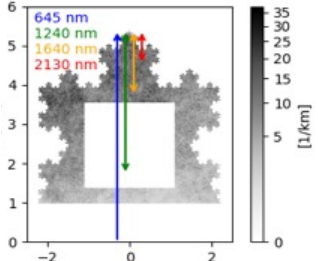


Diffusion process [#2]:

- random walks unfold in 3D *physical space*
- in the *veiled core*
- gradual loss of *positional memory*

Characteristic “diffusion scale,” L_d ,
i.e., the distance from sources
where it gets very dark:

$$L_d \times (\text{mean extinction}) = [3(1-\omega)(1-\omega g)]^{-1/2}$$



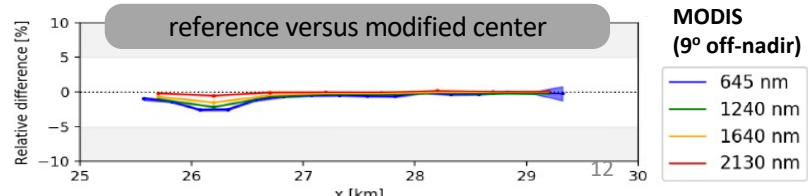
What happens to the now *close-to-isotropic* and already *somewhat-dispersed* forward- or backward-propagating solar radiation when it reaches the veiled core (VC)?

Let: H_{VC} = bulk size of VC; τ_{VC} = mean optical thickness of VC; and $\langle \rho^2 \rangle^{1/2}$ = RMS lateral transport along VC boundary, from entrance to escape. We know that for ...

- sensor on **illuminated side** [Davis et al., 1999ab]

$$\langle \rho^2 \rangle^{1/2} \sim H_{VC} / [(1-g) \tau_{VC}]^{1/2}$$
 → more opaque the VC, less the light will travel;
- sensor on **opposite side** [Davis & Marshak, 2002]

$$\langle \rho^2 \rangle^{1/2} \sim H_{VC}$$
 (irrespective of τ_{VC} and g)
 → light can escape from anywhere.



→ the VC grows with position across MODIS's SWIR channels

Cloud tomography forward model:

Need high accuracy ... *and efficiency!*



Diffusion process #1 & #2⁺ [#3⁺]:



- random walks unfold on 2D sphere (*direction space*)
- in the *outer shell*
- gradual loss of *directional* memory

➤ *standard 3D RT equation solver*

Diffusion process [#2]:

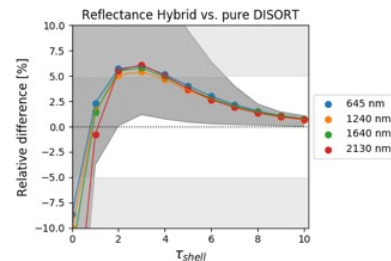
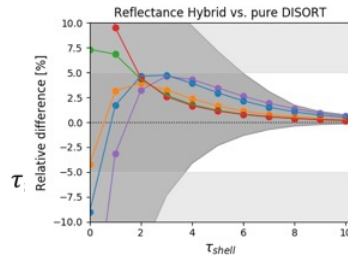
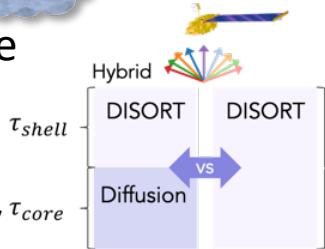


- random walks unfold in 3D *physical space*
- in the *veiled core*
- gradual loss of *positional* memory

➤ *efficient diffusion equation solver*

➔ Use best of both worlds in a **hybrid forward 3D RT model.**

Hybrid RT: Implementation in 1D



Hybrid RT: Possible Implementation in 3D

Propagation kernel (Beer's law): $T(\mathbf{x}, \Omega; s) = \exp[-\int_0^s \sigma_e(\mathbf{x} - \Omega s') ds']$ (BL)

$$I(\mathbf{x}, \Omega) = \int_0^{s_b(\mathbf{x}, \Omega)} [S(\mathbf{x}_s, \Omega) + Q(\mathbf{x}_s, \Omega)] T(\mathbf{x}, \Omega; s) ds$$

$$+ \begin{cases} 0 & \text{on } \partial M_{RT} \\ I_b(\mathbf{x}_b(\mathbf{x}, \Omega), \Omega) & \text{on } \partial M_{DA} \end{cases} T(\mathbf{x}, \Omega; s_b(\mathbf{x}, \Omega))$$
 (US)
$$S(\mathbf{x}, \Omega) = \sigma_s(\mathbf{x}) \int_{4\pi} p(\Omega \cdot \Omega') I(\mathbf{x}, \Omega') d\Omega'$$
 (SF)
$$Q(\mathbf{x}, \Omega) = F_0 T(\mathbf{x}, \Omega_0; s_b(\mathbf{x}, \Omega_0)) \sigma_s(\mathbf{x}) p(\Omega \cdot \Omega_0)$$
 (ST)

Boundary condition at $\mathbf{x}_b \in \partial M_{RT}$
 $I_b(\mathbf{x}_b, \Omega) = 0$, for $\mathbf{n}(\mathbf{x}_b) \cdot \Omega < 0$ (HBC)

Two-way coupling at $\mathbf{x}_b \in \partial M_{DA}$
RT → DA:
 $f_{in}(\mathbf{x}_b) = \int_{\mathbf{n}(\mathbf{x}_b) \cdot \Omega < 0} |\mathbf{n}(\mathbf{x}_b) \cdot \Omega| I(\mathbf{x}_b, \Omega) d\Omega$

DA → RT:
 $I_b(\mathbf{x}_b, \Omega) = f_{out}(\mathbf{x}_b) \frac{\mathbf{n}(\mathbf{x}_b) \cdot \Omega}{\pi}$, if $\mathbf{n}(\mathbf{x}_b) \cdot \Omega > 0$,
 where $f_{out}(\mathbf{x}_b) = \frac{1}{2} [1 + \chi \ell_t \mathbf{n}(\mathbf{x}) \cdot \nabla] J|_{\mathbf{x}=\mathbf{x}_b}$



Summary & Outlook

- 3D cloud tomography using multi-angle, multi-spectral, *and multi-pixel* data (i.e., images) collected from current and future **space-based sensors** remains a challenge.
 - Need *adapted* forward model (faster 3D RT solver)
 - Need *informed* inverse problem formulation/solution
 - ❖ **Definition of veiled core (VC) and its outer shell (OS) are key.**
- Deep dive into the physics of VNIR and SWIR cloud image formation, looking for insights ...
 - We uncover *two* complementary diffusion/random-walk processes:
 - ❖ First (in OS, near source) and last (in OS, near sensor) are **directional random walks on the 2D sphere** that end either in reflection or at the VC, with less and more dispersion, respectively.
 - ➔ **pixel-scale “features” → valid targets for detailed cloud tomography**
 - ❖ In the VC, solar radiation is transported by a standard **positional random walk in 3D space** that ends either in reflection or in transmission, with less and more dispersion, respectively.
 - ➔ **cloud-scale “R/T” contrast → only 1-or-2 unknowns for the whole VC**
- This applies to any passive observation of clouds in solar spectrum ... naked eyes included!

